# High-order Central Schemes for Hyperbolic Systems of Conservation Laws

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Joint work with E. Tadmor and C.C. Wu

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# Outline

- central schemes for hyperbolic conservation laws: overview and implementation
- ullet central schemes and MHD equations: the  $abla \cdot {f B} = 0$  constraint
- some examples: Euler equations of Gas Dynamics and Ideal MHD equations

# Hyperbolic Conservation Laws

We consider hyperbolic conservation laws in general

In one space dimension:

$$u_t + f(u)_{\times} = 0,$$

and two space dimensions:

$$u_t + f(u)_x + g(u)_y = 0,$$

with some initial data

$$u(x,y,0)=u_0(x,y),$$

where the Jacobian matrices  $\frac{\partial f}{\partial u}$  and  $\frac{\partial g}{\partial u}$  are diagonizable with real eigen values.

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   e.g., for MHD equations, we need to solve a large system with
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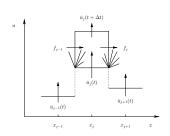
- onset of spurious oscillations
- additional challenges may come from the specific problem,
   e.g., for MHD equations, we need to solve a large system with an additional constraint
- we seek efficient numerical schemes capable of handling these challenges



# What do central schemes offer?

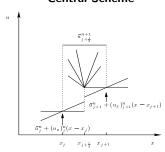
#### simplicity: no Riemann solvers

#### **Upwind Scheme**



requires a Riemann solver to distinguish from right- and left-going waves

#### Central Scheme



evolves solution over staggered grid, no Riemann solver is needed, but staggering requires smaller time step

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- highly adaptable implementation: minor changes required to solve different problems
- easy to parallelize: sequential function calls → concurrent function calls

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$$\frac{1}{\Delta x} \int_{t^n}^{t^{n+1}} \int_{x-\frac{\Delta x}{2}}^{x+\frac{\Delta x}{2}} u_t \ dt \ dx = -\frac{1}{\Delta x} \int_{t^n}^{t^{n+1}} \int_{x-\frac{\Delta x}{2}}^{x+\frac{\Delta x}{2}} f(u)_x \ dt \ dx$$

over the control volume  $[x_j, x_{j+\frac{1}{2}}] \times [t^n, t^{n+1}]$ ,

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over the control volume  $[x_j, x_{j+1}] \times [t^n, t^{n+1}]$ , this leads the equivalent cell average formulation

$$\bar{u}_{j+\frac{1}{2}}^{n+1} = \bar{u}_{j+\frac{1}{2}}^{n} - \frac{1}{\Delta x} \int_{t^{n}}^{t^{n+1}} \left[ f(u(x_{j+1}, t)) - f(u(x_{j}, t)) \right] dt$$

We now proceed in two steps:

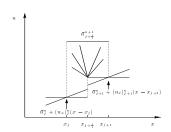
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We now proceed in two steps:



1. From the cell averages  $\{\bar{u}_{i}^{p}\}$ , a non-oscillatory polynomial reconstruction,

$$\tilde{u}(x,t^n) = \sum_{i} p_j(x,t^n) \cdot \mathbf{1}_{I_j},$$

is formed to recover  $\{\bar{u}_{j+\frac{1}{2}}^n\}$ ; where  $I_j = [x_j - \Delta x/2, x_j + \Delta x/2]$ .

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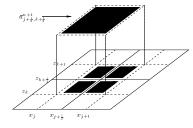
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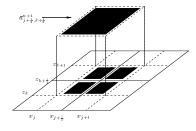
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corrector:

$$\bar{u}_{j+\frac{1}{2}}^{n+1} = \frac{1}{2} [\bar{u}_{j}^{n} + \bar{u}_{j+1}^{n}] + \frac{1}{8} [u_{j}' - u_{j+1}'] - \lambda [f(u_{j+1}^{n+\frac{1}{2}}) - f(u_{j}^{n+\frac{1}{2}})].$$



The staggered scheme can be extended to two space dimensions

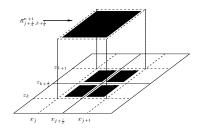


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$$u_{j,k}^{n+\frac{1}{2}} := \bar{u}_{j,k}^n - \frac{\lambda}{2} f_{j,k}' - \frac{\mu}{2} g_{j,k}',$$

where 
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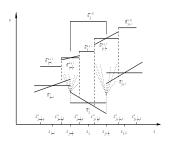
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corrector

$$\begin{split} \bar{u}_{j+\frac{1}{2},k+\frac{1}{2}}^{n+1} &= & \frac{1}{4}(\bar{u}_{j,k}^{n} + \bar{u}_{j+1,k}^{n} + \bar{u}_{j,k+1}^{n} + \bar{u}_{j+1,k+1}^{n}) + \frac{1}{16}(u_{j,k}' - u_{j+1,k}') \\ &- & \frac{\lambda}{2}\left[f(u_{j+1,k}^{n+\frac{1}{2}}) - f(u_{j,k}^{n+\frac{1}{2}})\right] + \frac{1}{16}(u_{j,k+1}' - u_{j+1,k+1}') - \frac{\lambda}{2}\left[f(u_{j+1,k+1}^{n+\frac{1}{2}}) - f(u_{j,k+1}^{n+\frac{1}{2}})\right] \\ &+ & \frac{1}{16}(u_{j,k}^{\lambda} - u_{j,k+1}^{\lambda}) - \frac{\mu}{2}\left[g(u_{j,k+1}^{n+\frac{1}{2}}) - g(u_{j,k}^{n+\frac{1}{2}})\right] \\ &+ & \frac{1}{16}(u_{j+1,k}' - u_{j+1,k+1}') - \frac{\mu}{2}\left[g(u_{j+1,k+1}^{n+\frac{1}{2}}) - g(u_{j+1,k}^{n+\frac{1}{2}})\right] \end{split}$$

## Modified central differencing (Kurganov and Tadmor, 2000)



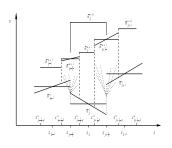
 Using the information provided by the local speed of propagation,

$$a_{j+\frac{1}{2}}^n = \max_{u \in \mathcal{C}(u_{j+\frac{1}{2}}^-, u_{j+\frac{1}{2}}^+)} \rho\left(\frac{\partial f}{\partial u}(u)\right),$$

where

$$u_{j+\frac{1}{2}}^+ := p_{j+1}(x_{j+\frac{1}{2}}) \ \ \text{and} \ \ u_{j+\frac{1}{2}}^- := p_{j}(x_{j+\frac{1}{2}}),$$

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we distinguish between the regions where the solution remains smooth –
 no Riemann fans, and regions where discontinuities propagate

- two sets of evolved values are calculated:
  - staggered values over non-smooth regions  $\{\bar{w}_{j+\frac{1}{2}}^{n+1}\}$
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- ullet or one can take the limit as  $\Delta t 
  ightarrow 0$  to arrive at the semi-discrete formulation:

$$\frac{d}{dt}\bar{u}_j(t) = \lim_{\Delta t \to 0} \frac{\bar{u}_j^{n+1} - \bar{u}_j^n}{\Delta t} = -\frac{H_{j+\frac{1}{2}}(t) - H_{j-\frac{1}{2}}(t)}{\Delta x},$$

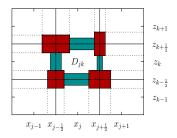
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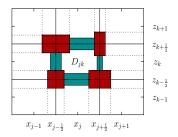
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$$\frac{d}{dt}\bar{u}_j(t)=-\frac{H_{j+\frac{1}{2}}(t)-H_{j-\frac{1}{2}}(t)}{\Delta x},$$

• where  $H_{j+\frac{1}{2}}(t) := \frac{f(u_{j+\frac{1}{2}}^+(t)) + f(u_{j+\frac{1}{2}}^-(t))}{2} - \frac{a_{j+\frac{1}{2}}(t)}{2} \left[ u_{j+\frac{1}{2}}^+(t) - u_{j+\frac{1}{2}}^-(t) \right]$ 

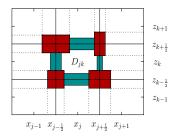


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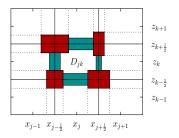
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staggered evolution over red cells



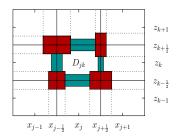
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- staggered evolution over red cells
- staggered evolution in one direction over green strips
- ullet non-staggered evolution over  $D_{j,k}$ , and

ullet reprojecting over original cells and taking the limit as  $\Delta t 
ightarrow 0$ , we arrive at:

$$\frac{d}{dt}\overline{u}_{j}(t) = -\frac{H_{j+\frac{1}{2},k}^{x}(t) - H_{j-\frac{1}{2},k}^{x}(t)}{\Delta x} - \frac{H_{j,k+\frac{1}{2}}^{z}(t) - H_{j,k-\frac{1}{2}}^{z}(t)}{\Delta z}$$

#### Central Schemes – Reconstruction

Example of a non-oscillatory reconstruction: second order *minmod* reconstruction (Van Leer, 1979)

$$p_{j,k}(x,z) = \bar{u}_{j,k}^n + u'_{j,k} \frac{(x-x_j)}{\Delta x} + u'_{j,k} \frac{(z-z_k)}{\Delta z}$$

where

$$u'(j,k) = minmod(\alpha \Delta_{+,x} \ \overline{u}_{j,k}^n, \frac{1}{2} \Delta_{0,x} \ \overline{u}_{j,k}^n, \alpha \Delta_{-,x} \ \overline{u}_{j,k}^n),$$

$$u'(j,k) = minmod(\alpha \Delta_{+,z} \ \overline{u}_{j,k}^n, \frac{1}{2} \Delta_{0,z} \ \overline{u}_{j,k}^n, \alpha \Delta_{-,z} \ \overline{u}_{j,k}^n),$$

with  $1 \leq \alpha < 2$ , and

$$minmod(x_1, x_2, \dots, x_n) = \begin{cases} min x_i, & \text{if} & x_i > 0 \ \forall i \\ max x_i, & \text{if} & x_i < 0 \ \forall i \\ 0 & \text{otherways} \end{cases}$$

#### Semi-discrete Central Schemes – Time Evolution

Solution evolved with SSP RK Schemes (Gottlieb et. al., 2001),

Example: Third-order scheme

$$u^{(1)} = u^{(0)} + \Delta t C[u^{(0)}],$$

$$u^{(2)} = u^{(1)} + \frac{\Delta t}{4}(-3C[u^{(0)}] + C[u^{(1)}]),$$

$$u^{n+1} := u^{(3)} = u^{(2)} + \frac{\Delta t}{12}(-C[u^{(0)}] - C[u^{(1)}] + 8 C[u^{(2)}]),$$

where

$$C[w(t)] = -\frac{H_{j+\frac{1}{2},k}^{x}(w(t)) - H_{j-\frac{1}{2},k}^{x}(w(t))}{\Delta x} - \frac{H_{j,k+\frac{1}{2}}^{z}(w(t)) - H_{j,k-\frac{1}{2}}^{z}(w(t))}{\Delta z}$$

# Ideal MHD Equations

conservation of mass:

$$\rho_t = -\nabla \cdot (\rho \mathbf{v}),$$

conservation of momentum:

$$(\rho \mathbf{v})_t = -\nabla \cdot [\rho \mathbf{v} \mathbf{v}^\top + (\rho + \frac{1}{2}B^2)\mathbb{I}_{3\times 3} - \mathbf{B}\mathbf{B}^\top],$$

conservation of energy:

$$e_t = -
abla \cdot \Big[ \Big( rac{\gamma}{\gamma - 1} 
ho + rac{1}{2} 
ho v^2 \Big) \; \mathbf{v} - \Big( \mathbf{v} imes \mathbf{B} \Big) imes \mathbf{B} \Big],$$

transport equation:

$$\mathbf{B}_t = \nabla \times (\mathbf{v} \times \mathbf{B})$$

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solenoidal constraint:

$$\nabla \cdot \frac{\partial \mathbf{B}}{\partial t} = \nabla \cdot [\nabla \times (\mathbf{v} \times \mathbf{B})] \quad \Rightarrow \quad \frac{\partial}{\partial t} (\nabla \cdot \mathbf{B}) = 0$$

## Solenoidal Constraint

• Numerical results indicate central schemes maintain  $\nabla \cdot \mathbf{B}$  small

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- Numerical results indicate central schemes maintain  $\nabla \cdot \mathbf{B}$  small
- Also, using the constraint transport approach and notation (Evans and Hawley, 1988), it can be shown that the magnetic field, as evolved by the second order fully-discrete staggered scheme (JT), can be written as

$$B_{j+\frac{1}{2},k+\frac{1}{2}}^{x,n+1} = \frac{1}{2} \left( b_{j,k+\frac{1}{2}}^{x,n+1} + b_{j+1,k+\frac{1}{2}}^{x,n+1} \right)$$

with

$$b_{j,k+\frac{1}{2}}^{x,n+1} = \tilde{b}_{j,k+\frac{1}{2}}^{x,n} - \frac{\Delta t}{\Delta z} \left( \Omega_{j,k+\frac{1}{2}}^{n+\frac{1}{2}} - \Omega_{j+1,k+\frac{1}{2}}^{n+\frac{1}{2}} \right),$$

and a similar expression for  $B^{z,n+1}_{j+\frac{1}{2},k+\frac{1}{2}}$ 



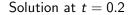
## MHD: Brio-Wu Rotated Shock Tube

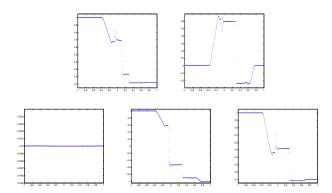
• One-dimensional Riemann problem with initial states given by

$$\left(\rho, \textit{v}_{\textit{x}}, \textit{v}_{\textit{y}}, \textit{v}_{\textit{z}}, \textit{B}_{\textit{x}}, \textit{B}_{\textit{y}}, \textit{B}_{\textit{z}}, \textit{p}\right)^{\top} = \left\{ \begin{array}{ll} \left(1, 0, 0, 0, 0.75, 0, 1, 1\right)^{\top} & \text{for } \textit{x} < 0 \\ \left(0.125, 0, 0, 0, 0.75, 0, -1, 0.1\right)^{\top} & \text{for } \textit{x} > 0 \end{array} \right.$$

- $\bullet$  Solved over a two dimensional domain with the direction of the flow rotated  $45^\circ$
- Solution computed up to t=0.2,  $x\in[-1,1]$ , with  $600\times600$  grid points,  $\gamma=2$ .

#### MHD: Brio-Wu Rotated Shock Tube





From top to bottom and from left to right: density, transverse velocity, transverse magnetic filed, parallel magnetic field, and pressure. The divergence of the reconstructed polynomial  $\sim 10^{-13}$ . Results computed with Jacobian free formulation of 2nd order JT scheme.

• Disruption of a high density cloud by a strong shock

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- Initial conditions

$$(\rho, v_x, v_y, v_z, B_x, B_y, B_z, \rho)^\top = \left\{ \begin{array}{ll} (3.86, 0, 0, 0, 0, -2.18, 2.18, 167.34)^\top & \text{for } x < 0.6 \\ (1, -11.25, 0, 0, 0, 0.564, 0.564, 1)^\top & \text{for } x > 0.6 \end{array} \right.$$

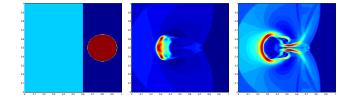
high density cloud –  $\rho$  = 10, p = 1 – centered at x = 0.8, y = 0.5, with radius 0.15,

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high density cloud –  $\rho$  = 10, p = 1 – centered at x = 0.8, y = 0.5, with radius 0.15,

• Solved up to t=0.06,  $(x,z)\in[0,1]\times[0,1]$ , with 256  $\times$  256 grid points, CFL number 0.5 and  $\gamma=5/3$ 



Solution of shock-cloud interaction, left: density at t=0, center: density at t=0.06, right: magnetic field lines at t=0.06. Results compputed with 3rd order semi-discrete scheme.

# Euler Equations of Gas Dynamics

conservation of mass:

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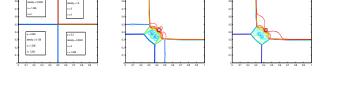
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equation of state:

$$p = (\gamma - 1) \left[ e - \frac{1}{2} \rho v^2 \right]$$

## Euler Equations: 2d Riemann Problem



Solution of a 2d Riemann problem, left: density at t=0 and initial conditions, center: density at t=0.3  $(S_{21}^{\leftarrow}, S_{32}^{\leftarrow}, S_{34}^{\leftarrow}, S_{41}^{\leftarrow}), \text{ right: pressure at t=0.3. Results compputed with 3rd order semi-discrete scheme using } 400 \times 400 \text{ grid cells.}$